

### MMGaiaDN Tech Workshop: Summary and Next Steps

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# Block 1: MWGaia Doctoral Network: Status and Updates

**Anthony Brown** 

- Network approx. at half-way point in terms of funding
- MWGaiaDN focused on Gaia Science and technical challenges for Gaia NIR
- Projects: MW (Mass-loss rates for RGB stars, kinematics, stell. characterisation, Exoplanet populations, Extinction Maps, L/SMG), extragal (Binary SMBH detection, Astrometric Quality of DR3 Quasars), Antone Investigation of rel. effects on high accuracy astrometry

Antonella Vallenari

- Publication/Outreach: Network presented at EAS 2024 & 2025
  - Science talks by multiple PhD candidates within MWGaiaDN

Despina Hatzidimitriou

Outreach Event planned close to Gaia DR4 release (schools across Europe)

Action: Incorporate GaiaNIR concept in to plans for 'Gaia Map' event planning

### Block 2: GaiaNIR: Overview and Status / Other existing and upcoming astrometry missions

#### Science Case Gaia NIR:

**David Hobbs** 

- Probe hidden regions. Up to 75% of MW stars could be mapped
- More precise proper motions by combining with Gaia precison ~ nas/yr
- Resetting the Gaia optical Reference Frame
- Low res. Dispersion spectra for most of the stars
- RV spectrograph for maybe 1 billion objects
- Exoplanets with periods up to 40 yrs
- Read-noise key in determining if GaiaNIR can outperform Gaia

#### JASMINE:

Ryouhei Kano

- Astrometry in gal. Nuclear region (1 kpc form centre / within 4kpc) --> 10<sup>4</sup> stars
- Transit obs. Around mid M type stars (aiming for earth-analogues)
- MDR passed in 2024/07 -> planned launch 2032
- Building interactions with the European community (Science, Data, Downlink).

## Block 2: GaiaNIR: Overview and Status / Other existing and upcoming astrometry missions

### Gaia's key impacts:

Action: Messaging: GaiaNIR gives the map of every component of the MW

### High-precision sky atlas

Enables detailed maps, including Kuiper Belt object outlines via occultation.

**Anthony Brown** 

### Accessible, vast dataset

Provides the community with easy access to an extensive and rich dataset.

### Dense astrophysical sampling

 Offers detailed CMD and phase space coverage (e.g., WD crystallization, Gaia-Enceladus merger, disc perturbations).

### Stellar and Galactic insights

 Delivers astrophysical parameters for many sources, informing studies of the Milky Way's oldest populations and disc dynamics.

### Survey calibration and time series

 Calibrates past/future surveys with a precise stellar reference frame; DR4 will include time series of astrometry, photometry, and spectra.

## Block 3: MWGaiaDN/ Industry Forum and Discussion

Luis Venancio

Gaizka Murga Llano

Joan-Manel Casalta

### Strong mission requirements

Ensure requirements are valid, clear, consistent, traceable, and realistically achievable.

### Science-industry collaboration

 Establish common requirement definitions and maintain good communication between scientists and industry.

### Lessons from past missions

**Rob Wilson** 

Learn from cases like Euclid where insufficient requirements led to issues (e.g., stray light, sunshield x-rays, focal plane size).

### Detector challenges for GaiaNIR

Minimize noise at high NIR gains, balance sensitivity with array size and TDI length, and keep design margins.
 Action: Inform/ engage with industry. Aim at TechWS#2 late 2026

#### Plan for evolution and risk

 GaiaNIR should not simply copy Gaia; allow for system evolution, mature key elements early, and have backups for high-risk components.

# Block 4: Science Requirement Input to Next Generation Astrometry Missions

David Hobbs

### Major technical challenges

Address crowding, high telemetry loads, detector performance (800–2300 nm), and read-noise constraints.
 Action: link with tools being developed in the MWGaiaDN

### Spectral and RV improvements

 Justify low-dispersion spectra, improve resolution, and plan for 1 billion radial velocities with slower scans, wider detectors, and lower noise.

#### Ambition for GaiaNIR

Aim for a large-scale GaiaNIR mission to extend capabilities, especially in near-infrared.

### Mock Catalogues

Marie Schölch

 Will help uncertainty estimates using modelling of 3D extinction, G mag etc., resulting in mock Gaia observations

#### Advance extinction studies

Alejandro Martín Escabia

 Use GaiaNIR to probe obscured regions and combine with optical data for more accurate extinction and 3D R(V) maps.

# Block 4: Science Requirement Input to Next Generation Astrometry Missions

Lessons from Gaia:

Action: review how lessons from Gaia might alter how re specifiy key requirements from GaiaNIR (e.g. BAV)

Robin Geyer

Ed Serpell

- Precise spacecraft tracking and synchronization
  - Ensure accurate determination of spacecraft orbit, velocity, and use of good clocks.
- Advanced calibration capability
  - Design systems to support complex calibration procedures and handle technical byproducts from solutions.
- Prepare for surprises
  - Maintain flexibility and readiness to address unexpected issues during the mission.
- Avoid disruptive mode changes
  - Minimize spacecraft mode switches (e.g., instruments, transmitters) to maintain stability.
- Design and testing insights
  - Understand and measure basic angle variations (BAV), isolate the launch adaptor, and account for multi-layer insulation (MLI) rigidity effects.

### Block 5: Near-IR astrometry missions: Current and Future technical studies

Detector Tests

Action: detector study outputs early 2026 → update GaiaNIR models

Nicholas Walton

 Mid-term study review Aug 2025. Currently Leonardo APDs looking very promising

### Metrology lessons from Theia/HWO

Fabien Malbet

- CMOS 46MP Gigapyx viable for high-precision astrometry
- Lab tests approaching spec; test bench operational
- Fringes & modulation measured; pixel calibration pending
- Optical correction ~10<sup>-2</sup> pixel level, aiming for 10<sup>-4</sup>
- GaiaNIR similar process, different detector materials

### Block 5: Near-IR astrometry missions: Current and Future technical studies

RV's for GaiaNIR:

Action: define RV window(s) and abundances/based on updated requirements -> WG activity

Szabolcs Mészáros

- Synthetic tests: 20,000 spectra with random Teff, log g, [M/H], vrad, SNR
- Optimal wavelength: 1910–2010 nm (strong Fe, Ca lines across parameter space)
- Backup options: 1430–1530 nm; other candidates between 1200–1600 nm
- Further analysis: 1 nm step, finer spectral sampling from 1200–2100 nm
- Additional potential: evaluate Teff, log g, [M/H], abundances in optimal range

### Deutsches Zentrum für Astrophysik

Martin Roth

- Large new lab for detector development
- Provide significant know-how for next generation ground/space based telescopes

### Block 6: Roadmap for Next Generation

Astrometry

Action: preparation for early proposal for mission selection

### Voyage 2050 timeline

L4 mission/concept CDF likely 2026 -> launch 2043

Nic Walton

- L5 science & mission development to commence 2027
  -> mission selection 2028/2029
- We need to ramp up communication efforts!

### Building the GaiaNIR community (<u>www.gaianir.org</u>)

David Hobbs

- GaiaNIR working groups need to be established/enlarged
- Consortium needs to be in place next year
- Co-I's must actively engage with the funding agencies and SPC representatives
- Codebase needs to be accessible and allow for development flexibility as well as adhering to ESA standards (learn from Vera Rubin etc.)
- Bluesky handle secured (gaianir.bsky.social) for the cost of one beer, thanks to Paul!

Action: website and communications campaign from q4/25

### **Next Steps**

Workshop brief report with list of actions

Thanks to all for their input and contributions!